

Stable, Bromine-Free, Tetragonal Perovskites with 1.7 eV Bandgaps via A-Site Cation Substitution

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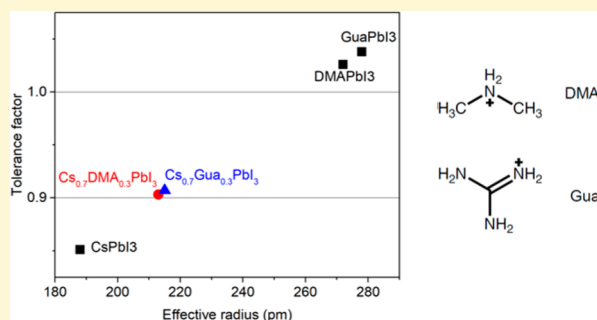


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Supporting Information

ABSTRACT: Metal halide perovskite absorbers with wide bandgaps (1.6–1.7 eV) that are suitable for tandem devices typically require high Br concentrations; this renders the material prone to halide segregation and degradation. Inorganic, bromine-free CsPbI₃ has a wide bandgap of 1.7 eV and does not suffer from halide segregation; however, these active layers are not stable at room temperature. Here, we report a method to create stable tetragonal perovskites with a bandgap near 1.7 eV: we add small amounts of large A-site cations having ionic radii between 272 and 278 pm—dimethylammonium (DMA) and guanidinium (Gua)—into the crystal lattice. When we deploy perovskites using mixed Cs and these large organic cations, we achieve stable, wide bandgap perovskites with power conversion efficiencies of 15.2% and V_{OC} of 1.19 V. This study extends materials selection for wide bandgap Cs-based perovskites.



Metal halide perovskites have attracted attention because of their application in low-cost, high-efficiency photovoltaics.¹ To make tandem solar cells with Si, a perovskite bandgap of 1.7 eV is optimal,² but usually requires compositions that contain 30–40% bromine (such as MAPbI_{0.6}Br_{0.4} and CsFAMAI_{0.6}Br_{0.4}).^{3–5} These compositions undergo I–Br phase segregation, lowering device stability and efficiency.⁶

To overcome halide segregation, it is desirable to minimize the Br concentration, but this lowers the bandgap to values that are suboptimal for tandem solar cells. Fully inorganic perovskites that use only iodine in their composition, such as CsPbI₃, can achieve wide bandgaps, but are not phase stable.⁷ For stable CsFA and CsMA mixtures, the ratio of Cs cannot exceed 30%, and the resulting compositions exhibit smaller bandgaps.⁸ Because of the small radius of the Cs⁺ ion, larger cations are needed to stabilize perovskite structures with Cs concentrations exceeding 30%. To maintain both a wide bandgap and phase stability in bromine-free perovskites, a lower concentration of larger organic cations is needed.

Here, we report bromine-free perovskites with bandgaps near 1.7 eV. We demonstrate that, by using A-site cations with ionic radii between 272 and 278 pm—DMA and Gua—we are able to create perovskites that are phase-stable at room

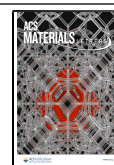
temperature. Compared to conventional cubic CsMAFA perovskites, we find the Cs_{1-x}DMA_xPbI₃ and Cs_{1-x}Gua_xPbI₃ perovskites exhibit tetragonal structures—a finding which we ascribe to the presence of the larger cation in the crystal lattice. In photovoltaic devices, we achieved a PCE of 15.2% using Cs_{1-x}DMA_xPbI₃, comparable to the PCE achieved in 40% Br wide bandgap perovskites. The operational stability of the Cs_{1-x}DMA_xPbI₃ devices is 230 h, nearly 20 times longer than that of 40% Br wide bandgap perovskites.

The phase instability of α -CsPbI₃ arises due to the small ionic radius of Cs⁺ cation, which results in a tolerance factor lower than that required for a stable perovskite phase.^{9,10} CsPbI₃ undergoes a phase transition to the non-perovskite δ -phase due to this instability, also accelerated by lattice distortions induced by polar solvents.^{7,11,12} Figure 1 and Table 1 show the tolerance factors of CsPbI₃, DMAPbI₃,

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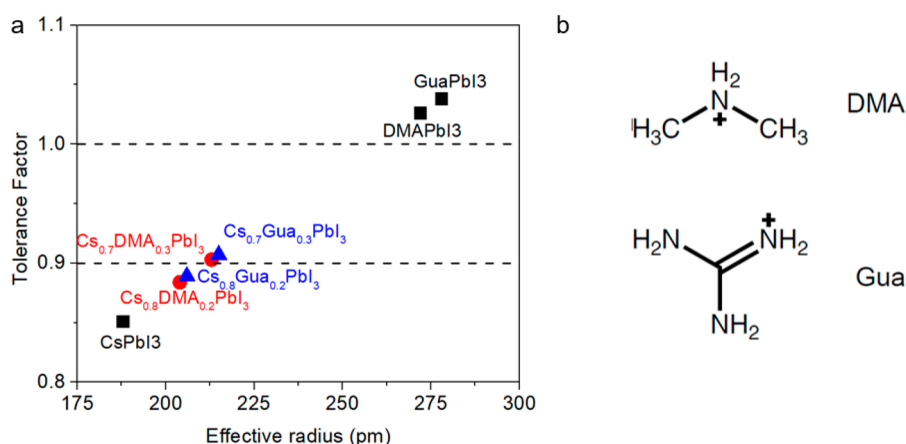


Figure 1. Tuning tolerance factor using mixed CsDMA and CsGua. (a) Tolerance factors of pure CsPbI₃, pure DMAPbI₃, pure GuaPbI₃, Cs_{1-x}DMA_xPbI₃ and Cs_{1-x}Gua_xPbI₃, (b) Molecular structure of DMA and Gua.

Table 1. Comparison of Ionic Radii for Different A-Site Cations

cation	ionic radius (pm)	tolerance factor for pure APbI ₃
DMA	272	1.026
Gua	278	1.038
MA	217	0.911
Cs	181	0.851

GuaPbI₃, and mixed CsDMA and CsGua perovskites. For pure CsPbI₃, the cubic phase gives a tolerance factor of 0.851, which is not in the range of structurally stable perovskites. Pure DMA and Gua cations are too large and give tolerance factors greater than 1; but mixed CsDMA and CsGua compositions

reach a tolerance factor slightly above 0.9 (Cs_{0.7}DMA_{0.3}PbI₃ gives 0.903 and Cs_{0.7}Gua_{0.3}PbI₃ gives 0.907), which falls within the favorable range for a stable perovskite phase.

We first examined the crystallinity and morphology of thin films of CsDMA and CsGua perovskites. Figure 2 a, b gives the morphology of the films under scanning electron microscope (SEM). The films exhibit smooth surfaces that are free of pinholes. Figure 2 c, d show X-ray diffraction (XRD) patterns of CsDMA and CsGua films. In a previous report, CsDMA perovskites synthesized at 100°C were assumed to be cubic,⁹ but the double peaks at 28° in that report suggested that the structure may actually be tetragonal. In this work, we anneal the film between 40 to 80°C, and the resulting XRD shows a tetragonal structure. Compared to normal tetragonal

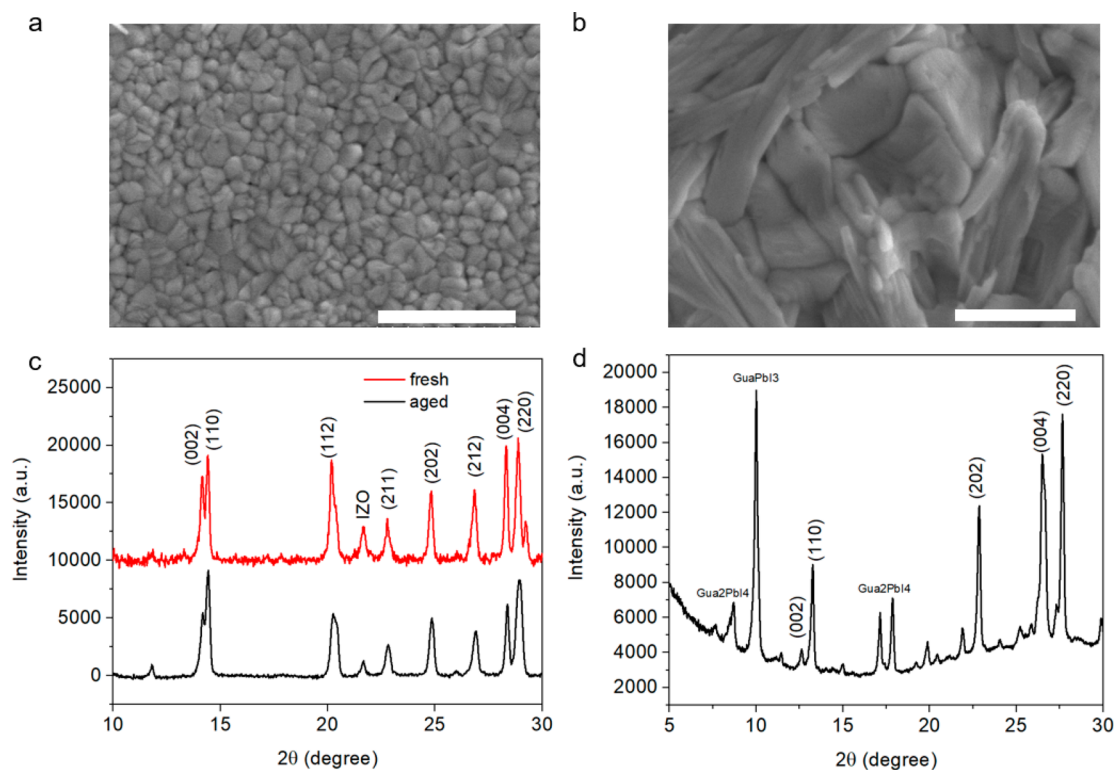


Figure 2. Morphology and crystallization of CsDMA and CsGua perovskite. (a) SEM images for CsDMA and (b) CsGua perovskite. The scale bar is 1 μm. (c) XRD pattern of fresh and aged CsDMA perovskite film and (d) XRD pattern of CsGua perovskite.

CsPbI₃ (β -CsPbI₃), pure β -CsPbI₃ is only observed above 210 °C, but for CsDMA and CsGua, the perovskite phase is formed between room temperature and 80 °C. This mixture method extends the range for stable, Cs-based perovskite. To confirm the presence of DMA/Gua in the annealed films, we carried out X-ray photoelectron spectroscopy (XPS) measurements (Figure S1). The N signal indicates that the large organic cations are present in the film following annealing at 80 °C.

Next, we monitored the stability of perovskite by obtaining XRD before and after ageing. Figure 2c gives a comparison of XRD patterns for the CsDMA perovskite. After samples age in air for 30 min, the XRD pattern of the CsDMA perovskite remains the same (Figure 2c), while in the case of pure CsPbI₃, it degrades considerably (Figure S2). We performed thermal gravimetric analysis (TGA) to assess thermal stability, and found no mass change below 200 °C (Figure S3), indicating that the CsDMA films are more thermally stable than perovskite compositions containing volatile MA molecules.¹³

After witnessing the ambient stability of the perovskite system, we proceeded to examine its optoelectronic properties. We collected steady state photoluminescence (PL) spectra and PL lifetimes of the CsDMA and CsGua perovskite films (Figure 3). The Cs_{0.7}DMA_{0.3} perovskite exhibits a PL emission

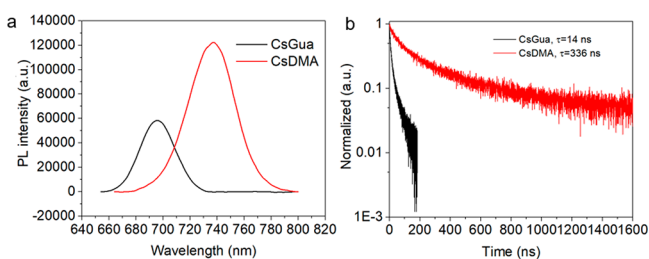


Figure 3. Photoluminescence spectroscopy of CsDMA and CsGua thin films: (a) Steady-state PL spectra and (b) PL lifetime of CsGua and CsDMA perovskite.

peak at 730 nm, corresponding to a bandgap at 1.70 eV, and the Cs_{0.7}Gua_{0.3} perovskite gives a PL emission peak at 690 nm, corresponding to a bandgap at 1.79 eV. This bandgap is similar to that of pure CsPbI₃ and is well-chosen for front cells in perovskite-Si (1.70 eV) and perovskite-perovskite tandems (1.79 eV). The PL lifetime of Cs_{0.7}DMA_{0.3} and Cs_{0.7}Gua_{0.3} perovskite are 336 ns and 14 ns. Compared to pure CsPbI₃, which gives a PL lifetime between 1 to 5 ns depending on the fabrication process,^{14,15} the mixed tetragonal perovskite gives a substantially longer PL lifetime comparable to the normal CsFA perovskite,¹⁶ suggesting reduced nonradiative recombination. We conclude that the mixed perovskite has a lower trap state density—a finding that should allow for a higher V_{oc} in perovskite solar cells (PSC). Pure CsPbI₃ PSCs usually suffer from a large V_{oc} loss: the highest reported V_{oc} is only 1.11 V, despite CsPbI₃ exhibiting a wide bandgap at 1.70 eV.

We, therefore, proceeded to investigate the effects of mixed tetragonal CsDMA and CsGua perovskites on the performance of solar cells. We fabricated planar PSCs with an architecture consisting of ITO/c-TiO₂/perovskite/Spiro-OMeTAD/Ag. The cross-sectional SEM image of the PSC is given in Figure S4. For Cs/DMA, we varied the Cs to DMA ratio (Table S1) and found the optimum power conversion efficiency (PCE) was achieved using 70% Cs and 30% DMA (Figure S5), corresponding to a bandgap of 1.70 eV. The $J-V$ curve of the champion device is given in Figure 4a. The champion device

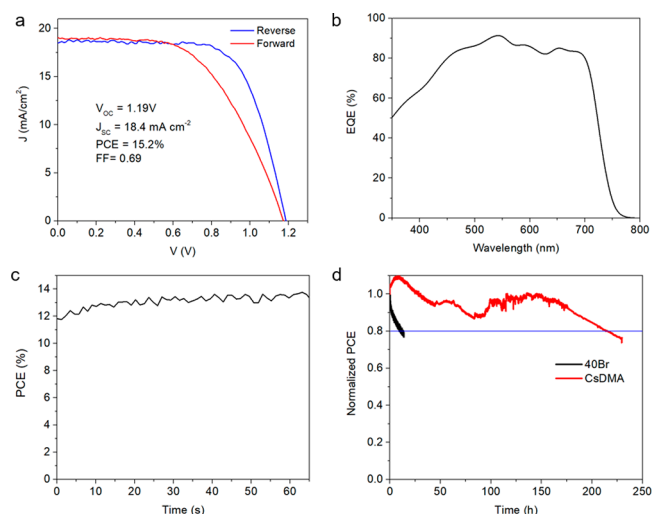


Figure 4. CsDMA device characterization. (a) $J-V$ characteristics of the champion device under simulated AM 1.5G solar illumination of 100 mW cm⁻². (b) EQE spectrum of the champion device. (c) Stable output at MPP, the champion device stabilizes at 13.5%. (d) Stability of an unencapsulated device under continuous white light LED illumination (100 mW cm⁻²) at MPP in N₂.

has a V_{oc} of 1.19 V, a J_{sc} of 18.4 mA cm⁻², and PCE of 15.2%. We achieved a V_{oc} higher than that reported in pure CsPbI₃ PSCs; it is 80 meV higher than in champion reports of CsPbI₃.¹⁴ Figure 4b gives the external quantum efficiency of the champion device, and it corresponds to an integrated photocurrent of 18.2 mA cm⁻². The stabilized output at the maximum power point (MPP) is 13.5% (Figure 4c). For CsGua, a similar optimization was carried out, but a maximum PCE of only 3.65% was achieved, possibly due to the impure phase of GuaPbI₃ and Gua₂PbI₄ (Figure S6).

We, then, checked the operational stability of the CsDMA perovskite PSCs. Since conventional wide-bandgap perovskites with Br content 30–40% are unstable due to the Hoke effect,¹⁷ we also fabricated PSCs using 40% Br perovskite active layers (Figure S7) using the same n-i-p structure to compare with the CsDMA devices herein.¹⁸ The 40% Br perovskite devices degrade rapidly to below 80% of initial PCE when operated at MPP operation for 12 h, while CsDMA devices are stable for 230 h under same conditions (Fig. 4d).

In summary, the present work reports stable, wide-bandgap, bromine-free perovskites. In contrast with CsMAFAPbI₃Br perovskite, we were able to tune the crystal lattice by incorporating even larger cations at a lower concentration. We demonstrate that by mixing CsPbI₃ with another A-site cation with an ionic radius between 272 and 278 pm—DMA or Gua—it is possible to form tetragonal perovskites that are stable at room temperature. Here, the large DMA or Gua cations are incorporated in the final lattice instead of being used as an additive. The mixed perovskite yields 14–60 times longer PL lifetimes than pure CsPbI₃, suggesting a lower trap state density. We achieved a PCE of 15.2% and a V_{oc} of 1.19 V. The V_{oc} of the CsDMA devices is higher than in prior reports of pure CsPbI₃ PSCs. The devices operated stably over 230 h at MPP, which is nearly 20 times longer than for the case of 40% Br wide-bandgap perovskites. This study extends the materials selection range for stable wide-bandgap perovskites.

■ ASSOCIATED CONTENT**SI Supporting Information**

The Supporting Information is available free of charge at <https://pubs.acs.org/doi/10.1021/acsmaterialslett.0c00166>.

Experimental details, XPS of N signal in Cs_{0.7}DMA_{0.3} perovskite, XRD of pure CsPbI₃ film after aging in air for 30 min, TGA of Cs_{0.7}DMA_{0.3}PbI₃, cross-sectional SEM image of CsDMA perovskite, performance of different CsDMA solar cells, and J–V plots of CsDMA solar cell with different DMA concentration, CsGua solar cell, and CsMAFA solar cell with 40% Br (PDF)

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Notes

The authors declare no competing financial interest.

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